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Sheffield.

# Do Massive Stars Form in Isolation?

Richard Parker<sup>1</sup>

Collaborator: Simon Goodwin<sup>1</sup>

1 - University of Sheffield, UK (email: [R.Parker@sheffield.ac.uk](mailto:R.Parker@sheffield.ac.uk))



# Introduction

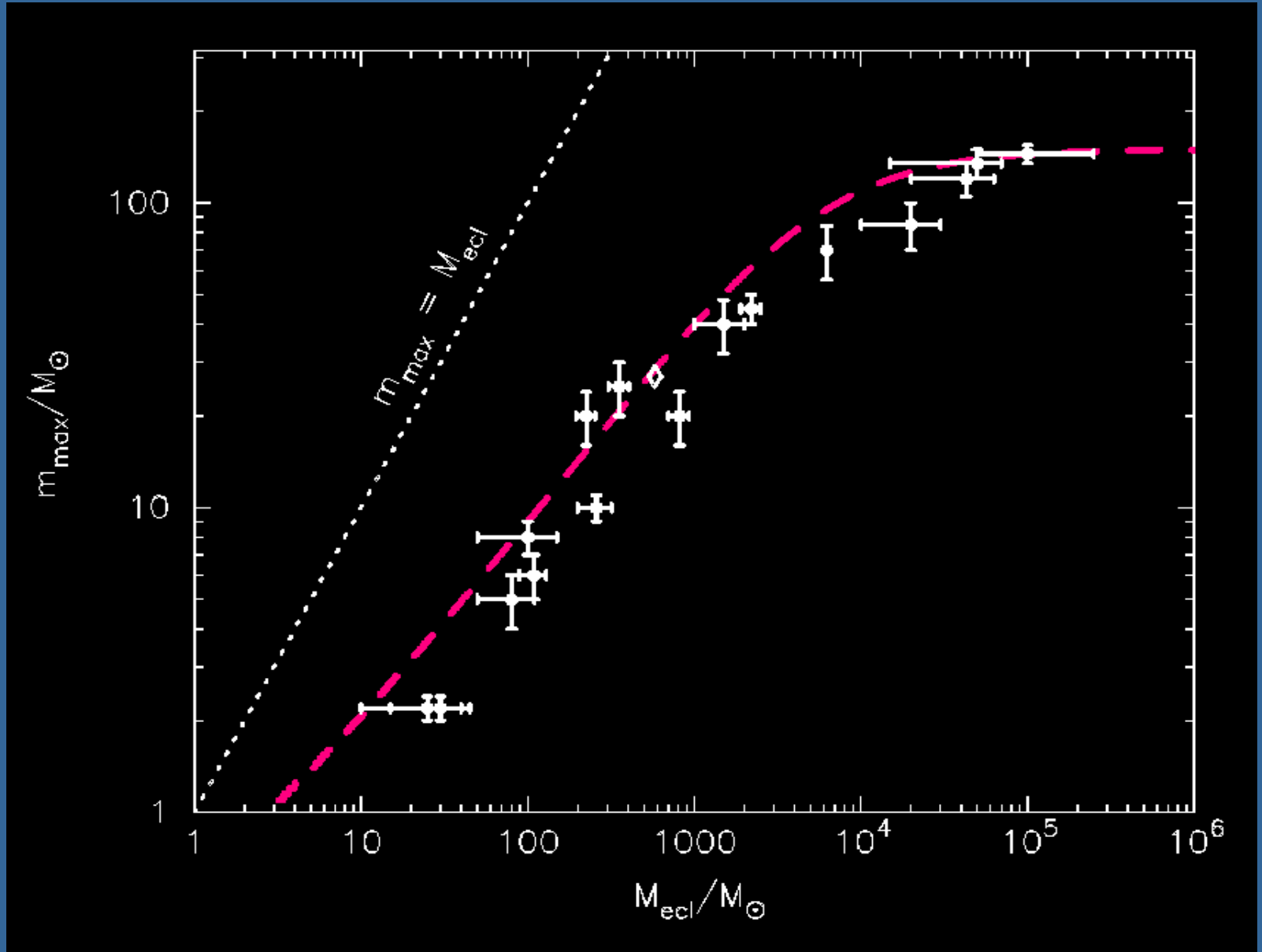
- de Wit et al. (2004) surveyed 43 O-type *field* stars.
- 12% found to be surrounded by a small cluster.
- Many were found to be runaways (see also Gvaramadze & Bomans 2008).
- $4(\pm 2)\%$  found to be in apparent isolation.
- 5% of B-type stars also observed in isolation.

# Introduction

- 70 – 90% of all stars form in clusters.
- Cluster masses described by single-power law slope;  $\beta = 1.5 - 2$ .
- Stars in these clusters appear to form with a universal IMF (Kroupa 2002).
- Therefore, one O-type star forms per 200 – 300  $M_{\odot}$  of stars.
- Is the mass of the most massive star in the cluster governed by the mass of the cluster (Weidner & Kroupa 2004, 2006; Weidner et al. 2009)?

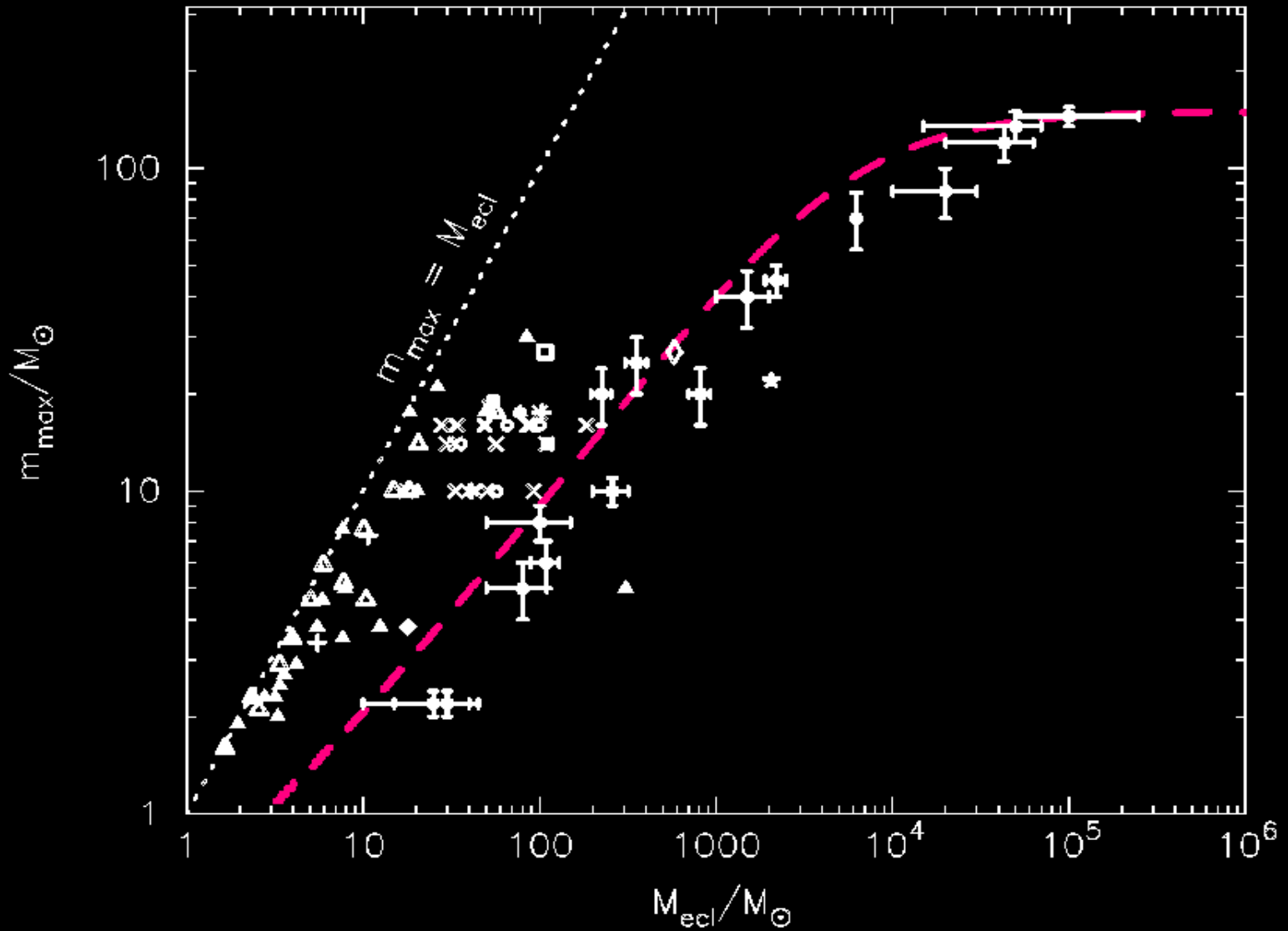


# The CMMSM





# Low-mass Clusters

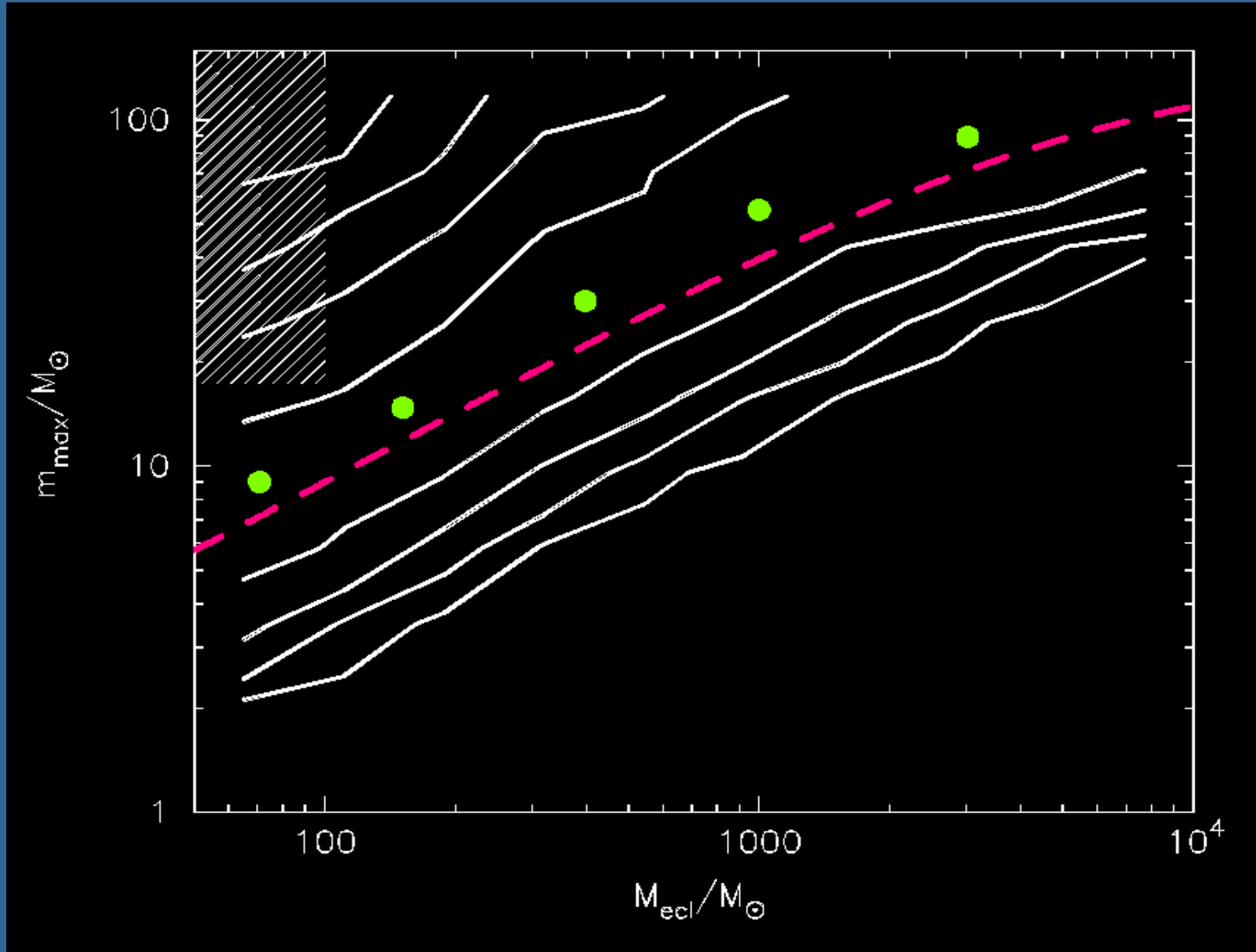


# Simulations

- Cluster masses chosen at random from CMF.
- Stellar masses:
  - i) randomly chosen from IMF
  - ii) constrained by a fundamental CMMSM
- 'Isolated' O-type star fraction ( $N(\text{sing.})/N_{\text{tot}}$ ):
  - i) 16% with no constraints
    - 5% if no B-type stars &  $M_{\text{ecl}} < 100 M_{\odot}$
  - ii) 4% with no constraints
    - 0% if no B-type stars &  $M_{\text{ecl}} < 100 M_{\odot}$

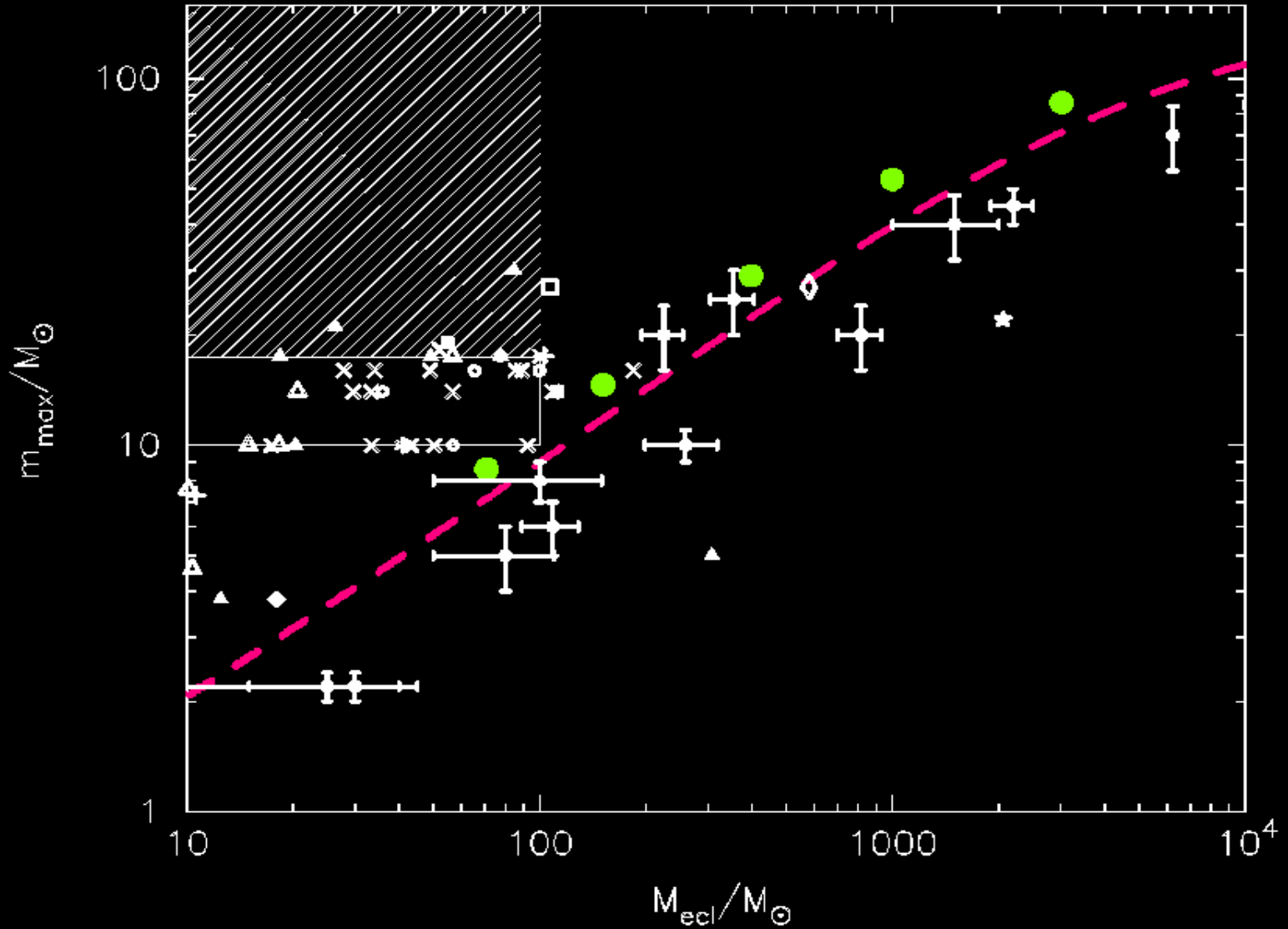


# A median CMMSM?





# A median CMMSM?



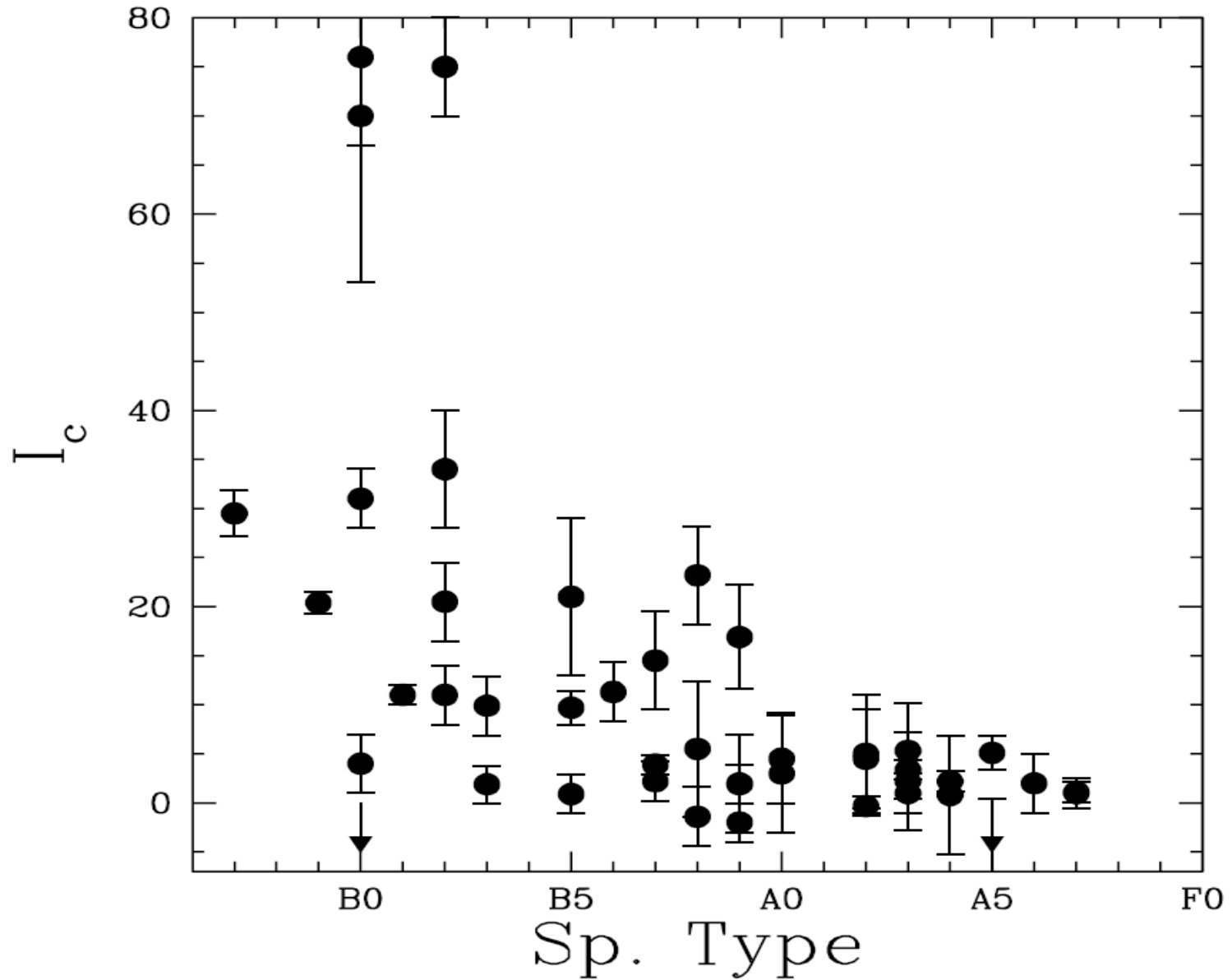


# Other Massive Stars

- In a series of papers, Testi et al. (1997, 1998, 1999) looked for evidence of clustering around Ae/Be stars.
- They plotted the spectral type of the most massive star in the cluster against the cluster 'richness indicator' – i.e. cluster mass.

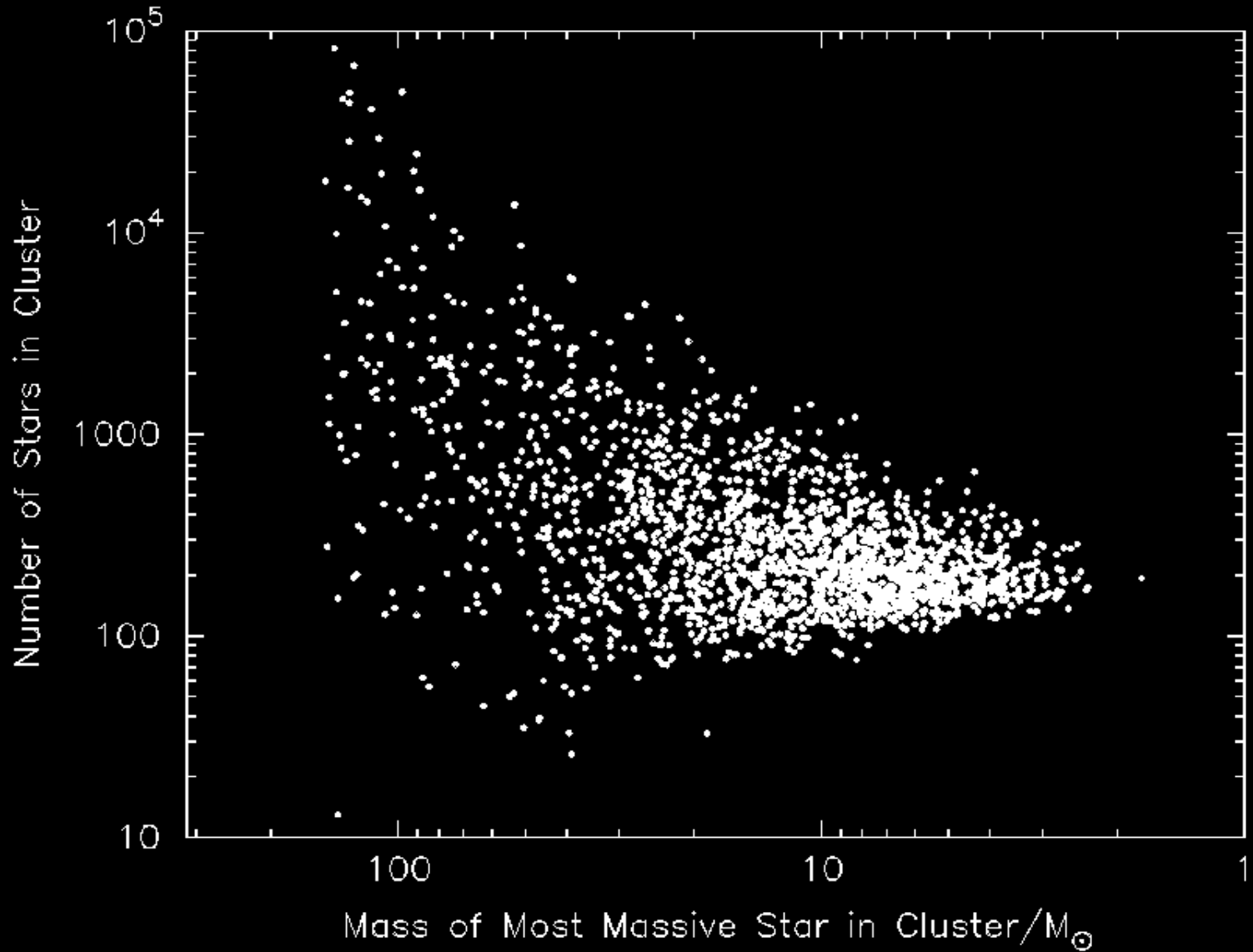


# Other Massive Stars



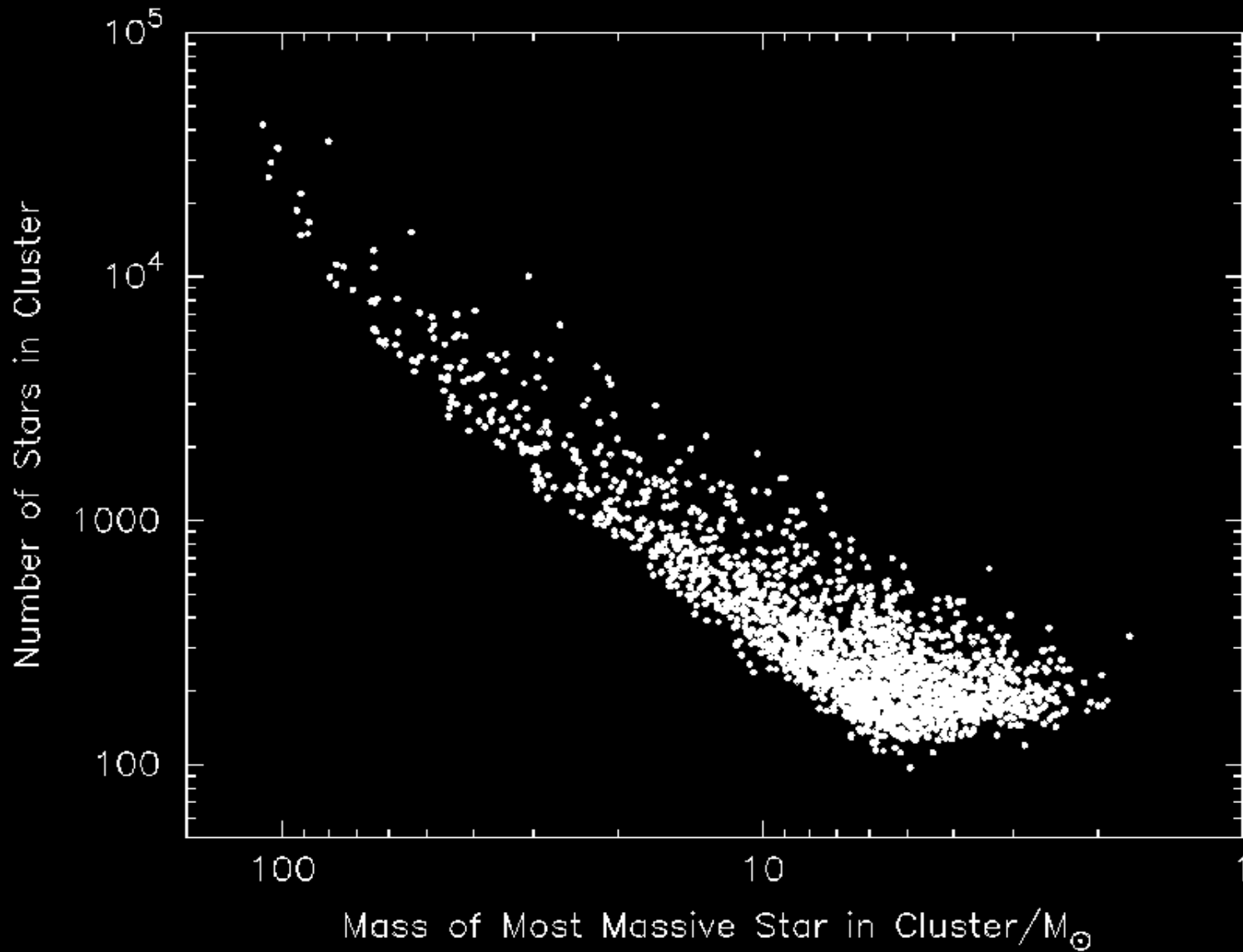


# Random Sampling





# CMMSM



# Conclusions

- $4 \pm 2\%$  field O-type stars apparently isolated.
- Random sampling produces 5% of low-mass clusters, with a single O-type star.
- Random sampling also recovers the *statistical* CMMSM relation.
- We argue against it being fundamental.

# References

- Parker & Goodwin (2007), MNRAS, 380, 1271
- Weidner et al. (2009), arXiv: 0909.1555
- Weidner & Kroupa (2006), MNRAS, 365, 1333
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- Maschberger & Clarke (2008), MNRAS, 391, 711
- de Wit et al. (2004), A&A, 425, 937
- Testi et al. (1999), A&A, 342, 515
- Testi et al. (1998), A&AS, 133, 81
- Testi et al. (1997), A&A, 320, 159