

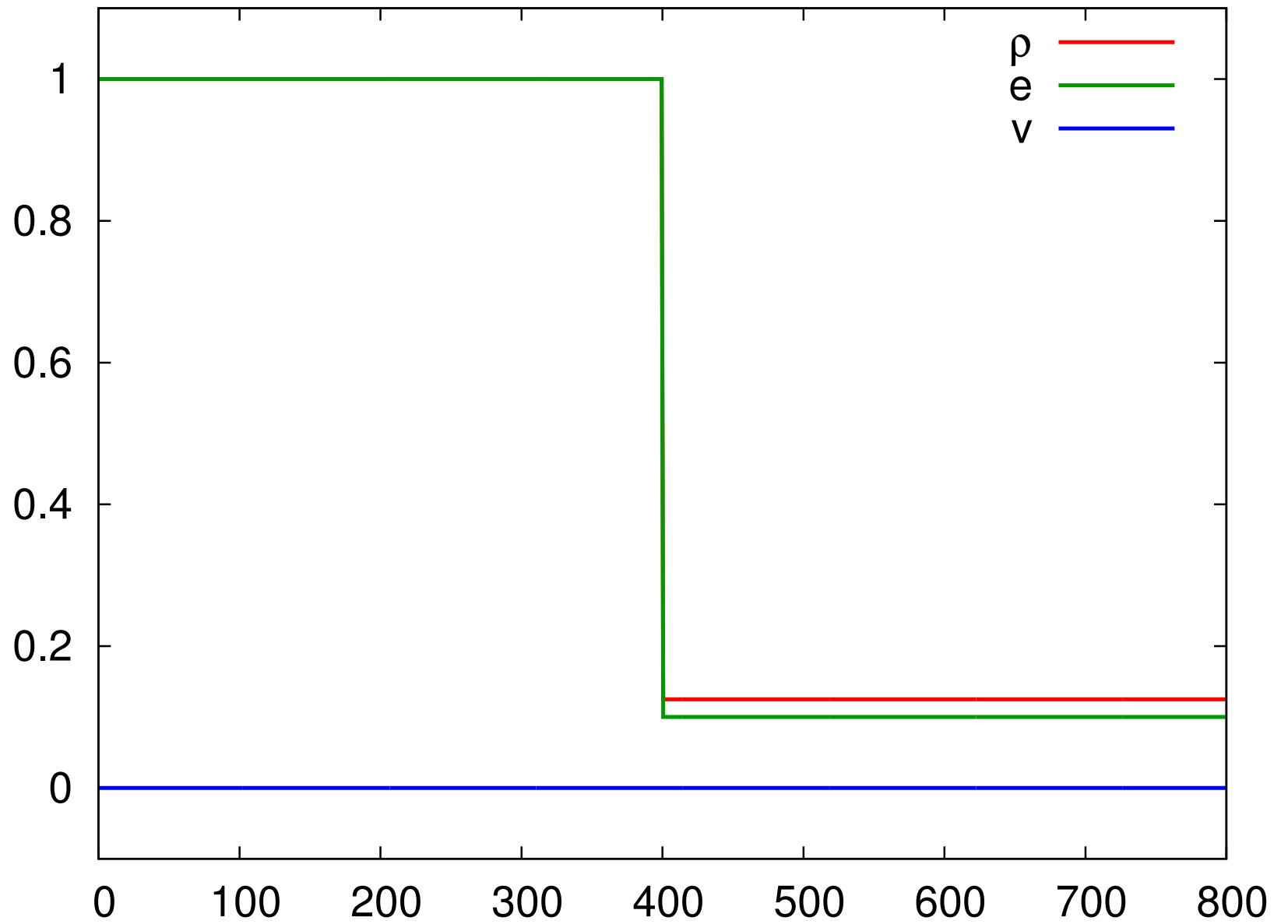
Hydrodynamic simulations of the ISM 2

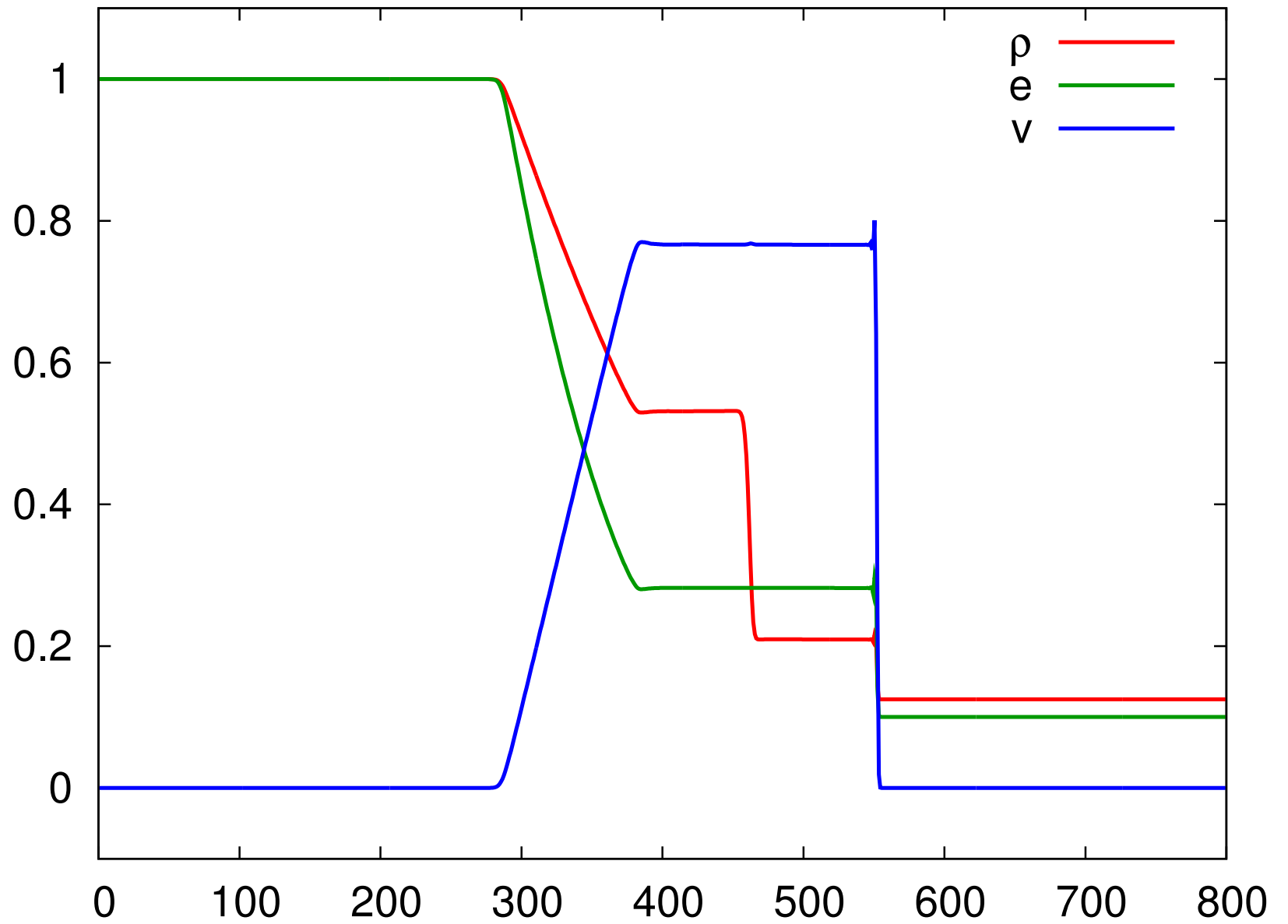
Richard Wunsch (richard@wunsch.cz)

Astronomický ústav AV ČR, Boční II 1401, Praha 4

Outline:

1. Basic numerical schemes:
Lax scheme, numerical viscosity, Courant-Friedrichs-Lewy condition
2. Introduction to ZEUS
3. Homework: Kelvin-Helmholtz instability with ZEUS





Finite Volume Methods - Conservative formulation

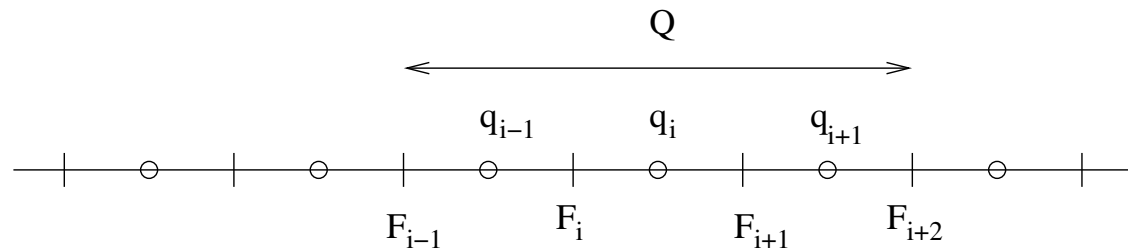
$$\frac{\partial}{\partial t} \begin{pmatrix} \rho \\ \rho v_i \\ e \end{pmatrix} + \frac{\partial}{\partial x_j} \begin{pmatrix} \rho v_j \\ \rho v_j v_i \\ e v_j \end{pmatrix} = \begin{pmatrix} 0 \\ -\frac{\partial p}{\partial x_i} - \rho \frac{\partial \Phi}{\partial x_i} \\ -\frac{\partial(\rho v_j)}{\partial x_j} - \rho v_j \frac{\partial \Phi}{\partial x_j} \end{pmatrix}$$

in general:

$$\frac{\partial q}{\partial t} + \frac{\partial F}{\partial x} = \Xi$$

Integrate over some volume, use divergence theorem:

$$\frac{\partial}{\partial t} \int_V q d^3V + \int_{\partial S} F d^2S = \int_V \Xi d^3V$$

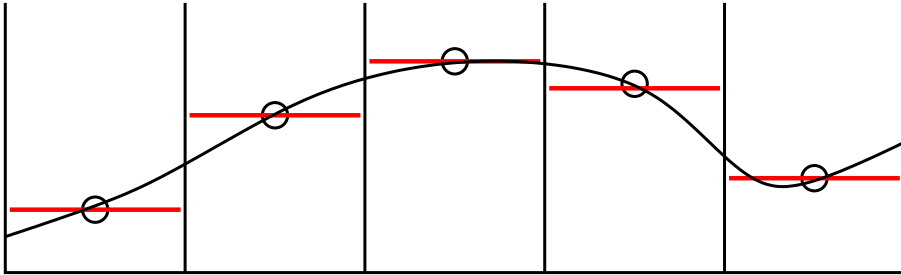


disregarding Ξ for a moment:

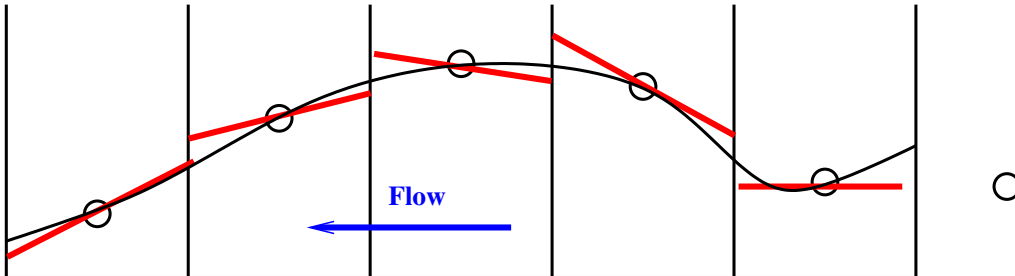
$$\frac{\partial q_i}{\partial t} = F_i - F_{i+1}, \text{ must work also for } Q : \frac{\partial Q}{\partial t} = \frac{\partial q_{i-1}}{\partial t} + \frac{\partial q_i}{\partial t} + \frac{\partial q_{i+1}}{\partial t} = F_{i-1} - F_{i+2}$$

Interpolation order

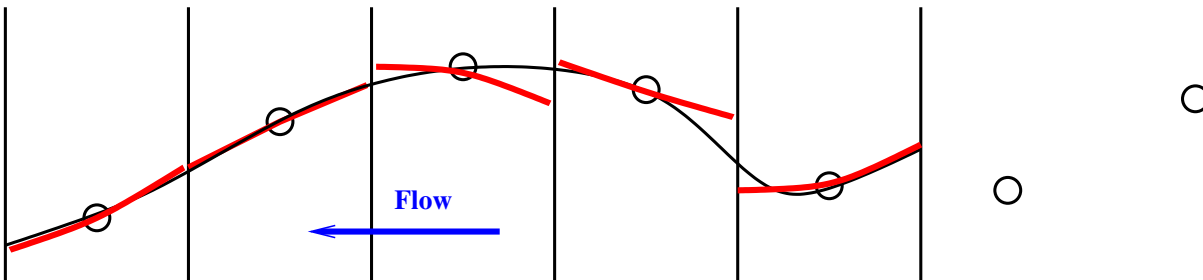
- 1st order, donor-cell



- 2nd order, piecewise linear

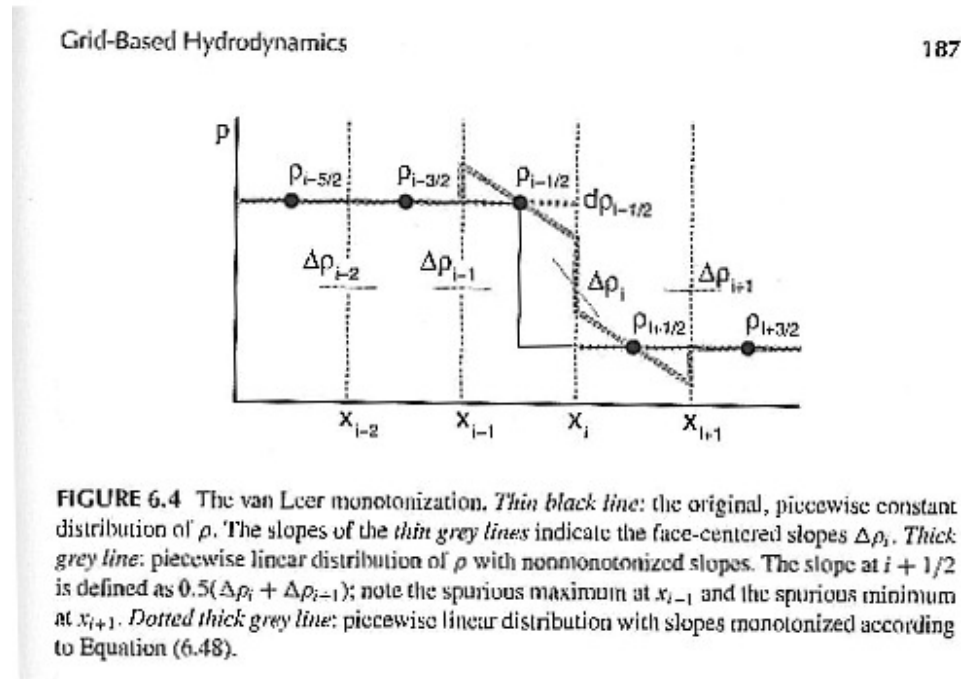


- 3rd order, piecewise parabolic advection (PPA)



Monotonicity

- van Leer (1977), to prevent spurious maxima and minima



- $\rho_i(w) = \rho_i + w d\rho_i$, where w is norm. distance from the cell centre
- the monotonicized slope $d\rho_i$ is:

$$d\rho_i = \begin{cases} \frac{2\Delta\rho_i\Delta\rho_{i+1}}{\Delta\rho_i + \Delta\rho_{i+1}} & \text{if } \Delta\rho_i\Delta\rho_{i+1} > 0 \\ 0 & \text{otherwise} \end{cases}$$

where $\Delta\rho_i = \rho_{i+1} - \rho_i$

Artificial viscosity

- necessary to dissipate a fraction of kinetic energy at shocks
- otherwise numerical instability appears
- von Neumann & Richtmeyer (1950):

$$\frac{v_i^{n+1} - v_i^n}{\Delta t} = -\frac{2(q_i^n - q_{i-1}^n)}{\Delta x(\rho_i^n + \rho_{i-1}^n)}$$
$$q_i = \begin{cases} C\rho_i(v_{i+1} - v_i)^2 & \text{if } (v_{i+1} - v_i) < 0 \\ 0 & \text{otherwise} \end{cases}$$

- sensitive only to compression
- satisfies conservation laws
- C - dimensionless constant, how many grid cells the shock will be spread out

ZEUS installation

- Mike Norman, Jim Stone, David Clarke
- official version 3.4.2 available at Laboratory for computational Astrophysics, University of California, San Diego
- D. Clarke version 3.5 (includes OpenMP)
- "easy" to install package:
`http://galaxy.ig.cas.cz/~richard/teaching/NAST021/`
- prerequisites:
 - ▷ UNIX type OS (tested with Linux)
 - ▷ Fortran77 and C compilers (tested with gfortran and gcc)
 - ▷ csh and ex (comes with vi text editor)
 - ▷ recommended: gnuplot, hdf4 libraries
- download `Zdistr-gfortran.tgz` and follow instructions in the README file
- compile the code and run the example `shkset_XYZ`
- use gnuplot script "`fig.gp`" to plot the results

The ZEUS code

- Operator splitting:

$$\frac{dy}{dt} = \mathcal{L}(y) = \mathcal{L}_1(y) + \mathcal{L}_2(y) + \dots$$

$$(y^1 - y^0) / \Delta t = \mathcal{L}_1(y^0)$$

$$(y^2 - y^1) / \Delta t = \mathcal{L}_2(y^1)$$

...

- Source step

$$\rho \frac{\partial v}{\partial t} = -\nabla P|_{SS1} - \rho \nabla \Phi|_{SS1} - \nabla \cdot Q|_{SS2}$$

$$\frac{\partial e}{\partial t} = -p \nabla \cdot v|_{SS3} - Q : \nabla v|_{SS2}$$

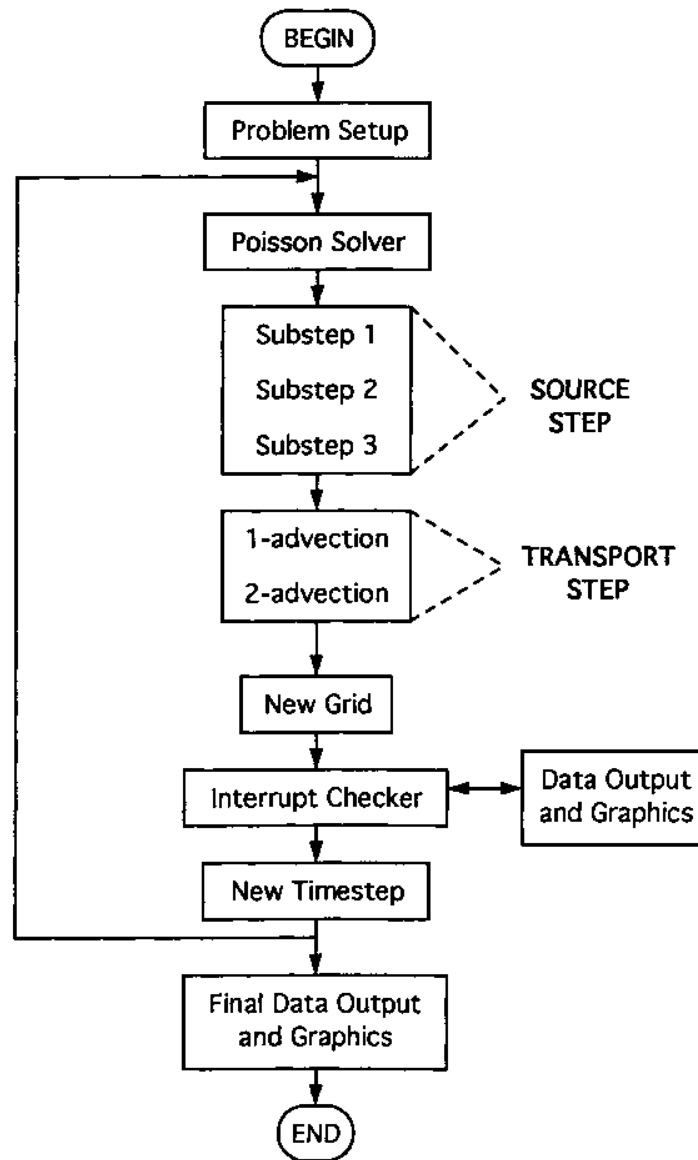
- Transport step

$$\frac{d}{dt} \int \rho dV = - \int_{\partial V} \rho (v - v_g) dS$$

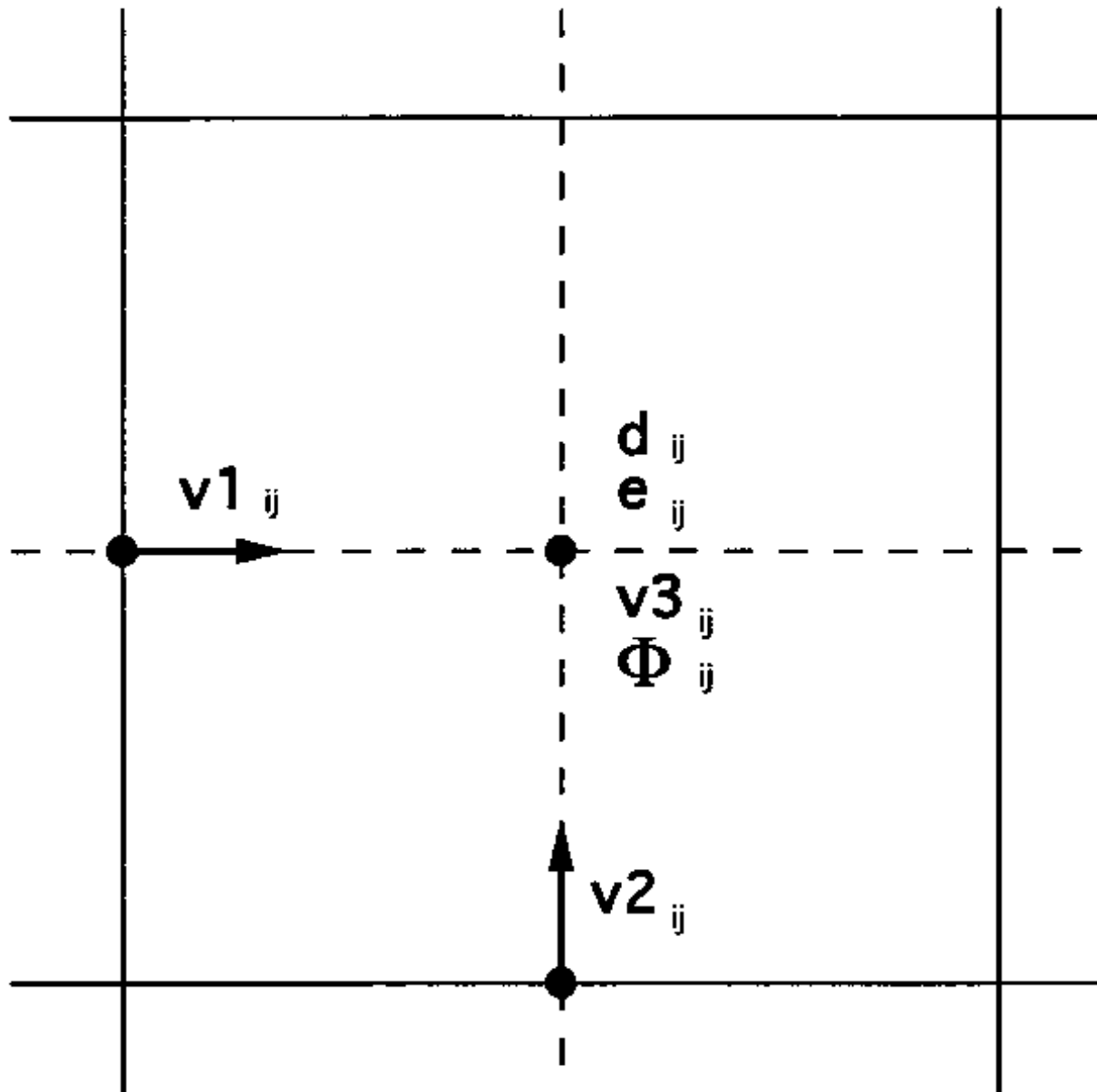
$$\frac{d}{dt} \int \rho v dV = - \int_{\partial V} \rho v (v - v_g) dS$$

$$\frac{d}{dt} \int e dV = - \int_{\partial V} e (v - v_g) dS$$

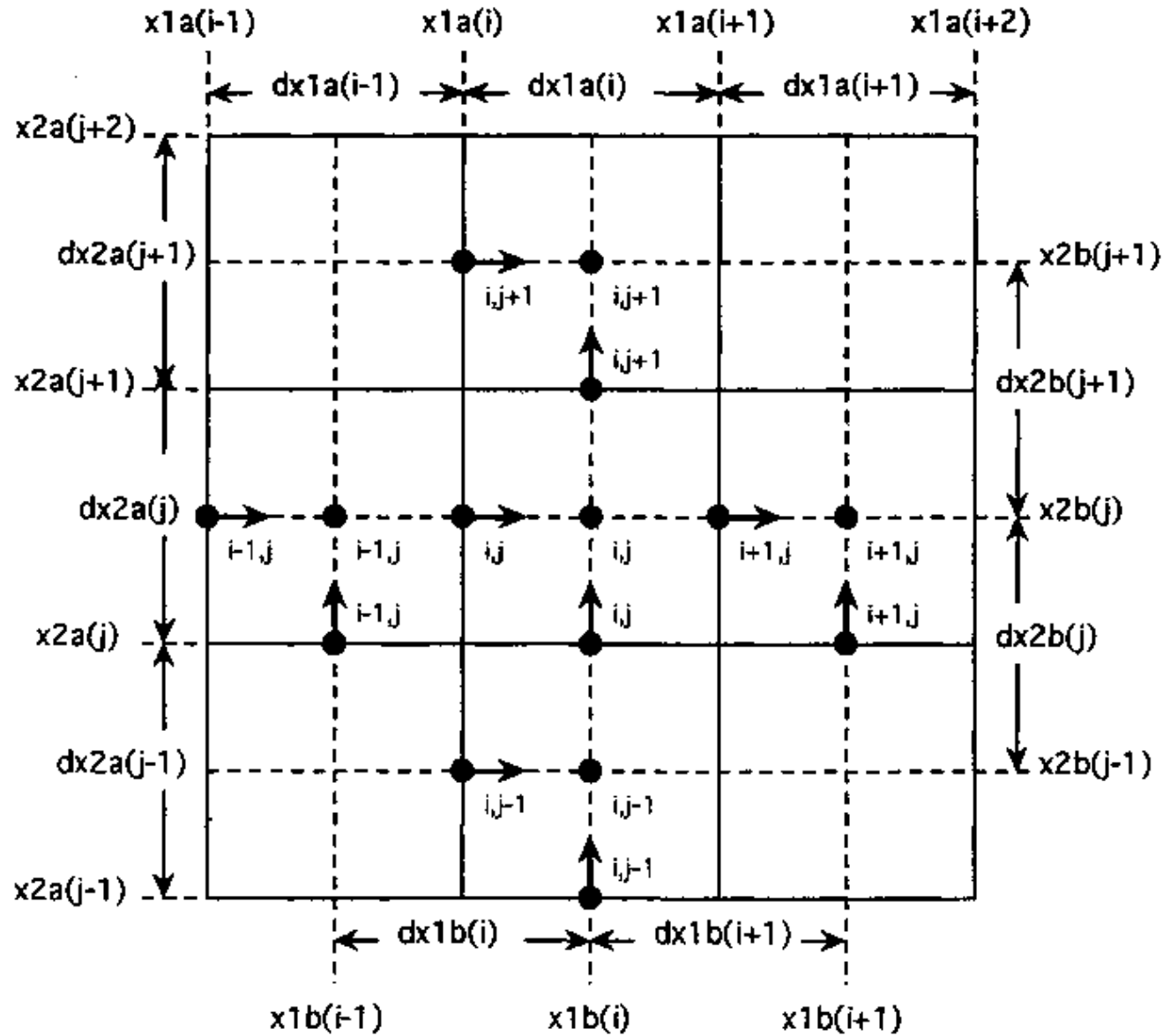
ZEUS - flowchart



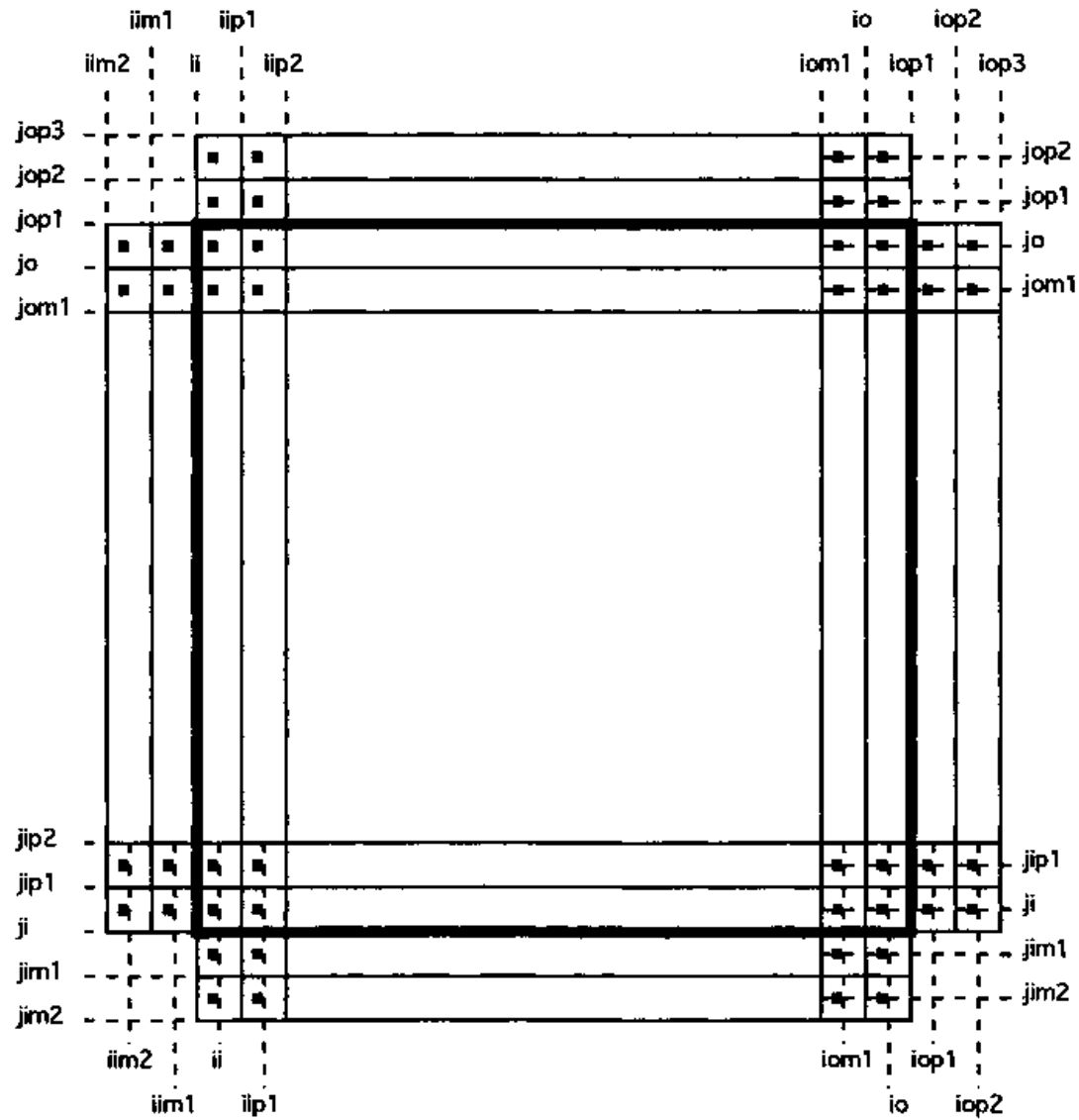
One ZEUS grid-cell



The ZEUS grid



Computational domain



ZEUS - How to setup a new problem

- Three steps: write a setup subroutine, compile and run
- Writing a setup subroutine
 - ▷ write it into the chengedeck in src/
 - ▷ see example chgz.cloud
 - ▷ read parameters from inzeus, set initial values on the grid, call bndyflgs and bndyall
- Compilation:
 - ▷ preprocessing by EDITOR (xedit22)
 - ▷ configured by the zeus34.mac file
 - ▷ zeus34.mac generated by zeus34.s script from the problem definition file in examples/
- Running:
 - ▷ ./xzeus34 reads parameters from inzeus
 - ▷ inzeus generated also by zeus34.s script from the problem definition file in examples/

You typically need to do

- write chgz.myproblem (in src/ directory)
- write problem definition file: examples/myproblem (and mention your chgz.myproblem in it)
- run commands:
cd myproblem
./~/ZEUS3D/arch/setZeus3D.Linux.gfortran
~/ZEUS3D/bin/zeus34.s ~/ZEUS3D/examples/myproblem
./xzeus34
- Try to run the simulation of Kelvin-Helmholtz instability in 2D (it will be equivalent of the exam for my part of this course)